1. Let  $S_n := \sum_{i=1}^n X_i$  be the empirical sum of iid random variables  $X_1, X_2, \ldots$  in  $\mathbb{R}$ . In the proof of the large-deviation principle for the empirical average  $\frac{1}{n}S_n$ , use Cramér's Theorem to show the alternative lower bound for the case where  $x < \mathbb{E}X_1$  and  $\epsilon > 0$  is sufficiently small, that is:

(LB') 
$$\liminf_{n \to \infty} \frac{1}{n} \log \mathbb{P}\left(\frac{1}{n} S_n \in B_{\epsilon}(x)\right) \ge -I(x).$$

2. For any two finite measures  $\rho, \nu \in \mathcal{M}(\mathcal{X})$  on a measurable space, the *relative entropy* is defined as

$$\mathcal{H}(\rho \mid \nu) := \begin{cases} \int_{\mathcal{X}} \left( \rho(dx) \log \frac{d\rho}{d\nu}(x) - \rho(dx) + \nu(dx) \right), & \rho \ll \nu, \\ \infty, & \text{otherwise} \end{cases}$$

Show that  $\inf_{\rho} \mathcal{H}(\rho \mid \nu) = 0.$ 

The next exercise does not require much preknowledge, but it shows a deep connection between large deviations, thermodynamics and information theory.

3. Consider *n* 'particles'  $x_1, \ldots, x_n$  on a finite space  $\{1, \ldots, L\}$ . We first study information theoretic aspects and later introduce randomness. In thermodynamics/statistical mechanics one often distinguishes between *microscopic states* and *macroscopic states*. In this setting, one may think of all particle coordinates  $x = (x_1, \ldots, x_n)$  as a micro state (where all particles are distinct), and its corresponding empirical measure  $\frac{1}{n} \sum_{i=1}^{n} \mathbb{1}_{x_i}$  as the macro state (where only the number of particles at each site matters), see picture:



(If you are having trouble with this exercise just take L = 2.)

(a) Write down the Boltzmann entropy of a given macro state  $\rho^n \in (\frac{1}{n}\mathbb{N})^L \cap \mathcal{P}(\{1,\ldots,L\})$ , defined as:

$$\operatorname{Ent}_{n}(\rho^{n}) := k_{B} \log \# \Omega_{n}(\rho^{n}), \quad \text{where}$$
$$\Omega_{n}(\rho^{n}) := \left\{ \text{micro states } x \in \{1, \dots, L\}^{n} : \text{macro state } \frac{1}{n} \sum_{i=1}^{n} \mathbb{1}_{x_{i}} = \rho^{n} \right\},$$

and  $k_B$  is the Boltzmann constant.

(b) The Boltzmann entropy blows up as  $n \to \infty$ , so for large particle numbers one usually takes the average entropy per particle. Calculate this limit (formally):

$$\operatorname{Ent}(\rho) := \lim_{n \to \infty} \frac{1}{n} \operatorname{Ent}_n(\rho^n),$$
  
assuming each  $\rho^n \in (\frac{1}{n} \mathbb{N})^L \cap \mathcal{P}(\{1, \dots, L\})$  and  $\rho^n \to \rho \in \mathcal{P}(\{1, \dots, L\}).$ 

*Remark*: the resulting formula is what physicists usually use as the entropy, whereas mathematicians usually drop the constant and flip the minus sign.

The Boltzmann entropy is purely combinatoric/information theoretic, so it is only a useful concept if all sites  $1, \ldots, L$  are equally likely. In general one needs something more advanced. Let  $X_1, X_2, \ldots$  be iid random variables with values in the finite set  $\{1, \ldots, L\}$ , with probabilities  $\mathbb{P}(X_1 = l) =: \nu_l$ . Now the *empirical measure*  $L^n := \frac{1}{n} \sum_{i=1}^n \mathbb{1}_{X_i}$  is a random variable in  $\mathcal{P}(\{1, \ldots, L\})$ .

- (c) For an abitrary test 'function'  $f \in \mathbb{R}^L$ , what is the weak limit of  $f \cdot L_n$  as  $n \to \infty$ ? No proof needed, but where does  $L^n$  weakly converge to?
- (d) Write down the probability that  $L^n = \rho^n$  for a given  $\rho^n \in (\frac{1}{n}\mathbb{N})^L \cap \mathcal{P}(\{1,\ldots,L\}).$
- (e) Again assuming  $\rho^n \to \rho \in \mathcal{P}(\{1, \dots, L\})$ , (formally) show that:

$$\lim_{n \to \infty} -\frac{1}{n} \log \mathbb{P}(L^n = \rho^n) = \mathcal{H}(\rho \mid \nu).$$

(f) In physics, one often has probabilities of the form

$$\nu_l = \frac{1}{Z_T} e^{-V_l/(k_B T)}$$

where  $V_l$  is some energy function, T is the temperature, and  $Z_T$  is a normalisation constant. Show that

$$\lim_{n \to \infty} -\frac{k_B T}{n} \log \mathbb{P}(L^n = \rho^n) = U(\rho) - T \text{Ent}(\rho) + \text{const}$$

for some function  $U(\rho)$  and a constant not depending on  $\rho$ .

*Remark*: physicists call  $U(\rho)$  the internal energy,  $U(\rho) - T\text{Ent}(\rho)$  the Helmholtz free energy, and they don't care about the constant.